

Nonlinear Schrödinger equation with random Gaussian input: Distribution of inverse scattering data and eigenvalues

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We calculate the Lyapunov exponent for the non-Hermitian Zakharov-Shabat eigenvalue problem corresponding to the attractive nonlinear Schrödinger equation with a Gaussian random pulse as an initial value function. Using an extension of the Thouless formula to non-Hermitian random operators, we calculate the corresponding average density of states. We also calculate the distribution of a set of scattering data of the Zakharov-Shabat operator that determine the asymptotics of the eigenfunctions. We analyze two cases, one with circularly symmetric complex Gaussian pulses and the other with real Gaussian pulses. We discuss the implications in the context of information transmission through nonlinear optical fibers.

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I. INTRODUCTION

One defining development in telecommunications technology during the last two decades has been the widespread use of optical fibers for transmitting enormous quantities of data across large—even transoceanic—distances. For such increasingly large distances, the nonlinearities in the fiber cannot be neglected, as they tend to distort transmitted pulses. Consequently, the detection of traditionally modulated signals becomes problematic. For fibers with negative group velocity dispersion (GVD) it is possible to compensate these effects by creating stable solitonic pulses [1,2]. As a first approximation, these solitary waves are solutions of the nonlinear Schrödinger equation (NLSE), the effective equation describing propagation of light in the frame comoving with the mean group velocity [3]. In normalized units the NLSE is expressed as

$$i\frac{\partial u}{\partial x} + \frac{\partial^2 u}{\partial t^2} + 2|u|^2u = 0, \quad (1)$$

where $u(t, x)$ is the (complex) envelope of the electric field, carrying the transmitted information signal along the fiber [4]. Traditional analyses of this equation focus on single and dilute solitonic propagation [5]. However, to address the ultimate information capacity limits through the fiber using solitonic pulses, one needs to explicitly consider dense soliton systems, where the soliton interactions can no longer be treated as small.

The problem of determining the spatial evolution of an incoming pulse $u(t) \equiv u(t, 0)$ is solved via the inverse scattering transform (IST), where $u(t)$ enters as the “potential” in a linear eigenvalue problem. For the NLSE this is the Zakharov-Shabat (ZS) eigenvalue problem [6], comprising of a 2×2 system of coupled first order differential equations,

$$\begin{pmatrix} i\partial_t & u^*(t) \\ -u(t) & -i\partial_t \end{pmatrix} \Psi_z(t) \equiv U(t)\Psi_z(t) = z\Psi_z(t), \quad (2)$$

where $\Psi_z(t) = [\psi_1(t)\psi_2(t)]^T$, and appropriate asymptotic conditions on the eigenstates, given in the next section.

In this paper we analyze the distribution of the scattering data, i.e., the average density of states (DOS) of U and the average distribution of a set of complex numbers $\{b_z\}$ that determine the asymptotics of the eigenstates, when $u(t)$ is drawn from a zero-mean, δ -correlated Gaussian distribution, describing the distribution of transmitted code words. Gaussian input signals are often used in information theory, and in linear transmission problems they often reach the Shannon capacity [7]. In addition, when the characteristic signal amplitude u_0 is much smaller than its bandwidth τ^{-1} (but with $D \equiv u_0^2\tau$ arbitrary), it is reasonable to approximate [8] the input distribution with a δ -correlated Gaussian for eigenvalues z small in the scale of τ^{-1} .

The non-Hermiticity of U causes the eigenvalues to spread over the complex plane. This generally makes the exact calculation of the DOS more difficult. Several powerful methods have been developed for calculating the statistical properties of non-Hermitian operators, which appear in the modeling of diverse physical processes (see, e.g., [9–21]). In most cases the random matrices are treated in a mean-field sense and are thus considered full random matrices. However, to our knowledge there are only a few non-Hermitian operators with diagonal randomness for which the exact density of states has been calculated in closed form [22–24]. In our case, we first calculate the Lyapunov exponent in closed form taking advantage of its self-averaging properties. Combining this with a generalization of the Thouless formula [25] for non-Hermitian operators [26], that relates the Lyapunov exponent with the DOS, we arrive at an explicit expression for the latter. Since the Lyapunov exponent is simply related to the localization length, it also provides information for the eigenfunctions of U .

In addition to the DOS we calculate the limiting distribution of the scattering data coefficients $\{b_z\}$, which depends strongly on the input distribution of $u(t)$: For circularly complex $u(t)$ the distribution of $\ln b_z$ approaches a Gaussian distribution albeit with singular variance growing as $T \ln T$, while for real $u(t)$ the distribution is highly singular, approaching a Cauchy distribution.

It should be noted that the Hermitian “counterpart” of this operator,

$$U_H = \begin{pmatrix} i\partial_t & u^*(t) \\ u(t) & -i\partial_t \end{pmatrix}, \quad (3)$$

arises in the IST for positive GVD, and also as a special case of the fluctuating gap model of disordered Peierls chains (see [27], and references therein). Its DOS and localization length have a long history of analysis [8,28–30].

The spectrum of U , together with the asymptotic behavior of the corresponding eigenstates Ψ_z , which as we shall see is determined by b_z , have the same information content as the input signal $u(t)$. This is because inverse scattering transform mapping between the scattering data of all eigenstates and $u(t)$ is one-to-one [31,32]. However, while the spatial evolution of $u(t, x > 0)$ and the eigenstates $\Psi(t, x > 0)$ is quite complicated, the eigenvalues z of U remain constant as the signal propagates down the fiber, and the corresponding scattering data vary in a trivial manner [31]. In fact, they can both be seen as playing the role of “action” variables changing adiabatically in the presence of nonintegrable perturbations. Therefore, the problem of light propagation in the fiber becomes easier to analyze in terms of the scattering data of the Zakharov-Shabat eigenproblem, especially in the presence of perturbations to Eq. (1), such as noise due to amplification or phase conjugation, which will ultimately determine the optical fiber capacity [33–36]. As a result, the description of the scattering data as a function of the input signal $u(t)$ may provide a framework for understanding the ultimate limits of information transfer through optical fibers.

II. LYAPUNOV EXPONENT AND DOS

We will now describe the basic steps to calculate the Lyapunov exponent of U in Eq. (2), which will then lead to the average DOS. To proceed, we start by introducing the ZS eigenvalue problem. Traditionally, this is defined as a scattering problem of the operator U in Eq. (2), in the presence of the potential $u(t) \equiv u(t, x=0)$, which decays sufficiently fast for $t \rightarrow \pm\infty$. In this context the scattering states are set up with the following asymptotic conditions outside the range of the potential:

$$\begin{aligned} \Psi_z(t) &\rightarrow \begin{pmatrix} 0 \\ 1 \end{pmatrix} e^{izt}, & \bar{\Psi}_z(t) &\rightarrow \begin{pmatrix} 1 \\ 0 \end{pmatrix} e^{-izt} & \text{as } t \rightarrow \infty, \\ \Phi_z(t) &\rightarrow \begin{pmatrix} 1 \\ 0 \end{pmatrix} e^{-izt}, & \bar{\Phi}_z(t) &\rightarrow \begin{pmatrix} 0 \\ 1 \end{pmatrix} e^{izt} & \text{as } t \rightarrow -\infty. \end{aligned} \quad (4)$$

For concreteness, we express the eigenvalue z as $z = \xi + i\eta$. The two sets of solutions in Eq. (4) are linearly related through the S matrix as follows:

$$\begin{bmatrix} \Phi_z(t) \\ \bar{\Phi}_z(t) \end{bmatrix} = \begin{pmatrix} b(z) & a(z) \\ \bar{a}(z) & \bar{b}(z) \end{pmatrix} \begin{bmatrix} \Psi_z(t) \\ \bar{\Psi}_z(t) \end{bmatrix}, \quad (5)$$

with the a 's and b 's being the transmission and reflection coefficients, respectively. By taking into account the symmetry of the problem under complex conjugation it is possible to show that $a(z^*) = \bar{a}^*(z)$ and $b(z^*) = \bar{b}^*(z)$, where the star (*) denotes the complex conjugate.

When the above solutions correspond to a localized eigenfunction with eigenvalue z , the transmission coefficient $a(z)$ has to vanish at that z , making the two sets of solutions directly proportional as follows:

$$\begin{aligned} \Phi_z(t) &= b_z \Psi_z(t), \\ \bar{\Phi}_z(t) &= -b_z^* \bar{\Psi}_z(t), \end{aligned} \quad (6)$$

where Φ_z and $\bar{\Phi}_z$ are the admissible exponentially decaying eigenfunctions for $\text{Im}(z) > 0$ and $\text{Im}(z) < 0$, respectively. Note that inside the region where $u(t)$ is finite, they should decay with a Lyapunov exponent $\kappa(z)$, rather than with $|\text{Im}(z)|$ as in Eq. (4). The proportionality constants b_z in Eq. (6) are not simply related to the functions $b(z)$ evaluated at the eigenvalue z [37]. It is clear from above that delocalized states can only exist when $\text{Im}(z) = 0$.

The proportionality factors b_z and their corresponding eigenvalues z are very important quantities in the theory of the inverse scattering transforms: Together with the continuum delocalized states characterized by $b(z)$, they can completely reconstruct the original $u(t)$. Therefore, in the context of information theory, they carry the same information content. In physical terms, the localized eigenstates of the Zakharov-Shabat problem correspond (through the IST) to the solitonic excitations in the fiber, while the continuous spectrum for $\text{Im}(z) = 0$ gives the radiation modes, which spread out and decrease in amplitude as the signal propagates down the optical channel. We will focus on the localized states, since in the limit $T \rightarrow \infty$ they correspond to the dominant part of the solution.

Our computation of the DOS of the problem is based on the calculation of the Lyapunov exponent $\kappa(\xi, \eta)$, which then yields the density of states through the generalized Thouless formula (derived in the Appendix) as follows:

$$\rho(\xi, \eta) = \frac{1}{2\pi} \left(\frac{\partial^2}{\partial \xi^2} + \frac{\partial^2}{\partial \eta^2} \right) \kappa(\xi, \eta). \quad (7)$$

The (upper) Lyapunov exponent is defined by

$$\kappa = \lim_{t \rightarrow \infty} \frac{1}{2t} \ln[|\psi_1(t)|^2 + |\psi_2(t)|^2], \quad (8)$$

which can also be written as

$$\kappa = \lim_{t \rightarrow \infty} \frac{1}{2t} \int_0^t dt' \frac{\partial}{\partial t'} \ln(|\psi_1(t')|^2 + |\psi_2(t')|^2). \quad (9)$$

Since the system is self-averaging (the evolution of ψ_1, ψ_2 along t is a Markov process), we can exchange the average over t in Eq. (9) with an average over the Gaussian ensemble as follows:

$$\kappa = \frac{1}{2} \lim_{t \rightarrow \infty} \left\langle \frac{\partial}{\partial t} \ln(|\psi_1|^2 + |\psi_2|^2) \right\rangle. \quad (10)$$

This is our starting point for calculating κ . From Eq. (2) we find:

$$\partial_t(|\psi_1|^2 + |\psi_2|^2) = 2\eta(|\psi_1|^2 - |\psi_2|^2). \quad (11)$$

Defining the complex variable $f(t) = \frac{\psi_1(t)}{\psi_2(t)} = e^{w(t)+i\phi(t)}$, with $w \in (-\infty, \infty)$ and $\phi \in [0, 2\pi)$, we can rewrite Eq. (11) as

$$\partial_t \ln(|\psi_1|^2 + |\psi_2|^2) = 2\eta \tanh w. \quad (12)$$

We are interested in the long-time behavior of Eq. (12). For a given $u(t)$, $w(t)$ undergoes constant change at any t , but the probability distribution of its values in the Gaussian ensemble will tend to a stationary distribution for large t . To see this, we must derive the Fokker-Planck equation for the joint probability distribution $P(w, \phi; t)$. This is straightforward for δ -correlated Gaussian potentials, since in this case $w(t)$ and $\phi(t)$ become Markov processes [38].

A. Circular complex Gaussian potential $u(t)$

We start by calculating the density of states (DOS) $\rho(\xi, \eta)$ and localization length $l(\xi, \eta)$ when $u(t)$ is circularly symmetric, i.e., $u(t) = \frac{1}{\sqrt{2}}[u_1(t) + iu_2(t)]$, with u_1, u_2 real, $\langle u_i(t) \rangle = 0$, and $\langle u_i(t)u_j(t') \rangle = D\delta_{ij}\delta(t-t')$, $i, j = 1, 2$. In this case, the evolution of w and ϕ is described by the set of stochastic equations as follows:

$$\begin{aligned} \partial_t w &= 2\eta + 2 \cosh w (\sin \phi u_1 + \cos \phi u_2), \\ \partial_t \phi &= -2\xi - 2 \sinh w (\cos \phi u_1 - \sin \phi u_2). \end{aligned} \quad (13)$$

The Fokker-Planck equation derived from these (in the Stratonovich picture) is

$$\partial_t P = \partial_v [(1-v^2)(-2\eta + D\partial_v)]P + \partial_\phi \left(2\xi + D \frac{v^2}{1-v^2} \partial_\phi \right) P, \quad (14)$$

where $v = \tanh w$. A simplification can be obtained by integrating over ϕ . Because the right-hand side of Eq. (12) depends only on v , we only need $\tilde{P}(v) = \int_0^{2\pi} d\phi P(v, \phi)$ to calculate the average. Integrating over ϕ and using the periodicity of P in this variable we find the Fokker-Planck equation for \tilde{P} as follows:

$$\partial_t \tilde{P} = \partial_v [(1-v^2)(-2\eta + D\partial_v)]\tilde{P}. \quad (15)$$

Setting the left-hand side to zero we find the stationary solution to which the system relaxes for large t as follows:

$$\tilde{P}(v) = \frac{\eta e^{2\eta v/D}}{D \sinh \frac{2\eta}{D}}. \quad (16)$$

This is also a stationary solution of the full Fokker-Planck equation (15), implying that asymptotically ϕ becomes uniformly distributed. We can now calculate the Lyapunov exponent from Eqs. (10) and (12) as follows:

$$\kappa = \eta \int_{-1}^1 dv v \tilde{P}(v) = \frac{D}{2} \left[\frac{2\eta}{D} \coth \left(\frac{2\eta}{D} \right) - 1 \right]. \quad (17)$$

Note that for large $|\eta|$, $\kappa \approx |\eta|$ independently of D : this is expected since in this limit the potential decouples the left (ψ_1) from the right moving (ψ_2) wave functions. A simple

application of the Thouless formula (7), gives the exact density of eigenstates for the system as follows:

$$\rho(\xi, \eta) = \frac{2}{\pi D} \frac{\frac{2\eta}{D} \coth \left(\frac{2\eta}{D} \right) - 1}{\sinh^2 \left(\frac{2\eta}{D} \right)}. \quad (18)$$

The independence of ρ from ξ is not surprising: the density of states of the Hermitian (diagonal) part of Eq. (2) is independent of ξ . Therefore, in the so-called mean-field approximation [12,39] the extension in the imaginary axis will be ξ -independent. It should be noted, however, that that mean-field approach would have given a *constant* DOS within a zone around $\eta=0$, rather than Eq. (18). A comparison of this expression with the result of numerical simulations can be seen in Fig. 1. Again note that for large η the density of states vanishes: in this limit there is an exponentially small probability for finding a potential deep enough to create a bound state.

The localization length $l(\xi, \eta)$ is the inverse of the Lyapunov exponent, $l = \kappa^{-1}$. To see this, we note that the Wronskian of two independent solutions of Eq. (2) is constant, therefore if for a given z it has a solution increasing exponentially as $\exp(\kappa t)$, its other solution has to be exponentially decreasing as $\exp(-\kappa t)$. Thus a square integrable solution necessarily decays with length scale κ^{-1} inside the support of $u(t)$. From Eq. (17) one can see that states become increasingly delocalized as the eigenvalues approach the real axis on the complex z plane: l diverges as $l \sim \frac{3D}{2\eta^2}$ near the real axis. The localization length also determines the stability of the corresponding eigenvalue to the presence of a finite time window of the pulse T . Specifically, the typical lifetime of a state with eigenvalue z will scale as $\sim e^{\kappa(\eta)T/2}$ [8]. Indeed we see this in Fig. 1, where for states with localization length comparable to the system length T , i.e., close to $\eta=0$, the calculated DOS is no longer valid. To capture the behavior of the DOS in this region, a zero-dimensional analysis similar to [21,39,41,42] is needed.

B. Real initial pulse $u(t)$

We can also analyze the opposite case when $u(t)$ is real, Gaussian with $\langle u(t) \rangle = 0$, and $\langle u(t)u(t') \rangle = D\delta(t-t')$. In this case, the evolution of w and ϕ is described by Eq. (13) by setting $u_2(t) = \text{Im}(u(t)) = 0$. The corresponding steady state solution of the Fokker-Planck equation can be derived from

$$\begin{aligned} 0 = \partial_v [(1-v^2)(\sin^2 \phi D\partial_v - 2\eta)]P + 2\xi \partial_\phi P \\ - \sin 2\phi \partial_\phi \left[\frac{1+v^2}{1-v^2} + 2v\partial_v \right] P + \frac{2v^2 \cos^2 \phi}{1-v^2} \partial_\phi^2 P. \end{aligned} \quad (19)$$

For large $|\xi|$, P is independent of ϕ to leading order in ξ . Therefore the large- ξ expansion is essentially identical to a Fourier expansion. Integrating Eq. (19) over ϕ gives Eq. (15). Thus \tilde{P} is to leading order identical to that of the circularly symmetric complex u . After some algebra one can derive the next-leading-order result. To order $\mathcal{O}(1/\xi^2)$ the correction to the Lyapunov exponent is

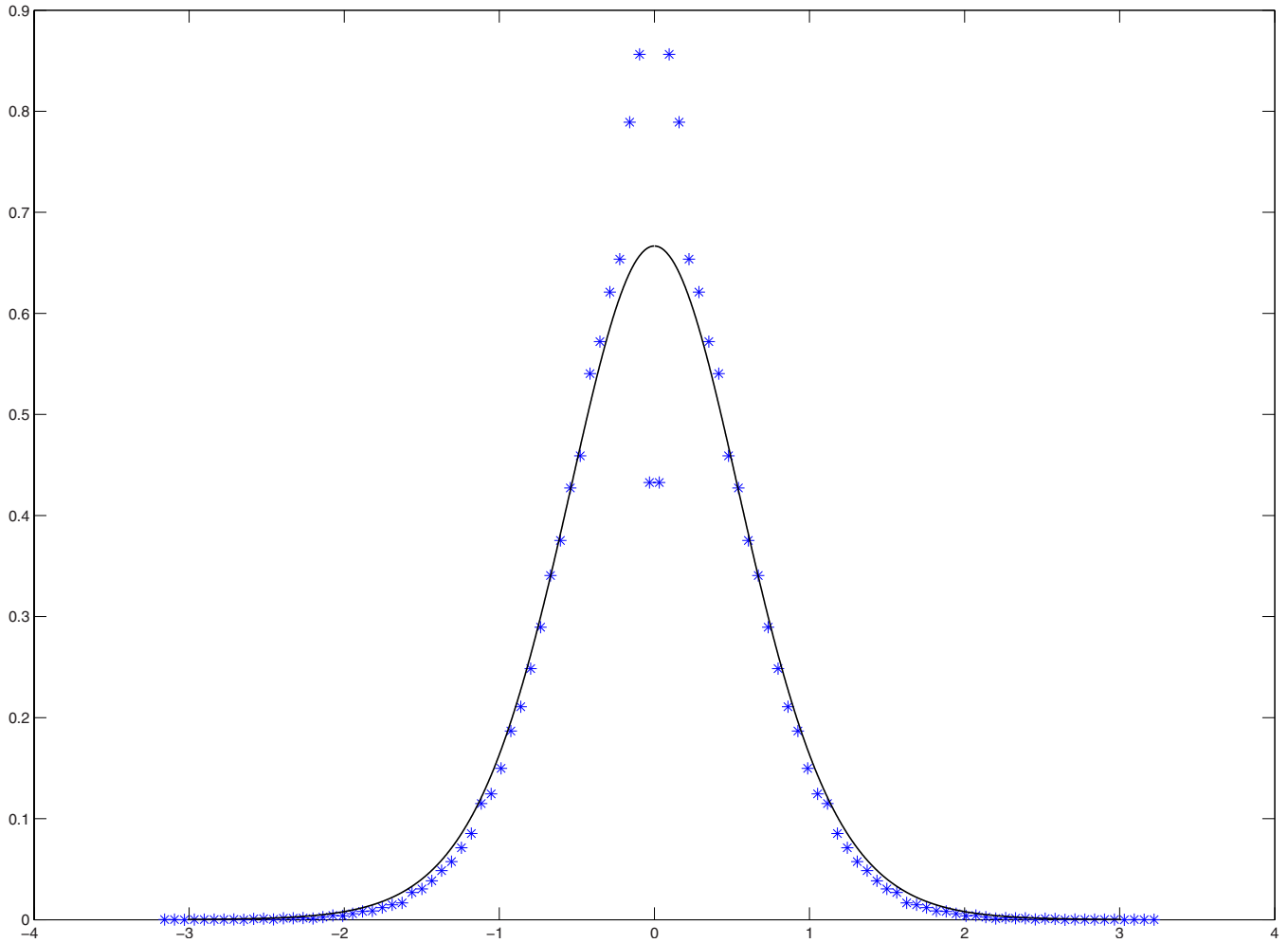


FIG. 1. (Color online) Theoretical curve (solid line) and results of numerical simulations for the profile of the DOS vs η . We have used the modified Ablowitz-Ladik diagonalization scheme [40] to ensure that the determinant of the discretized transfer matrix has unit value. The value of D is 1, the size of the system is $T=135$, and the step size is 0.075. The disturbance near $\eta=0$ is a finite-size effect. The localization length grows as $l \sim 3D/2\eta^2$ near $\eta=0$ and so numerical results differ from our $T \rightarrow \infty$ formula for $|\eta| \lesssim \sqrt{\frac{D}{T}}$.

$$\delta\kappa = \frac{D^2}{4\xi^2} \left[1 - \frac{\eta}{D} \coth\left(\frac{2\eta}{D}\right) + \frac{2\eta^2}{D^2 \sinh^2\left(\frac{2\eta}{D}\right)} \right], \quad (20)$$

$$\kappa(\eta) = \frac{\eta I_1(\eta/D)}{I_0(\eta/D)}, \quad (23)$$

resulting in the following correction to the DOS expression of Eq. (18):

$$\delta\rho = \frac{D^2}{\pi\xi^2} \frac{3\eta \coth\left(\frac{2\eta}{D}\right) - \frac{6\eta^2}{D \sinh^2\left(\frac{2\eta}{D}\right)} - \frac{4\eta^2}{D}}{\sinh^2\left(\frac{2\eta}{D}\right)}. \quad (21)$$

In the opposite limit of small ξ , we expect the distribution in ϕ to be peaked. Indeed for $\xi=0$, Eq. (19) has a solution that is proportional to $\delta(\cos \phi)$. This results in

$$\tilde{P}(v) = \frac{e^{\eta v/D}}{\pi I_0(\eta/D) \sqrt{1-v^2}}, \quad (22)$$

with corresponding Lyapunov exponent

where $I_{0,1}$ are modified Bessel functions of the first kind. We see that compared to Eq. (16), (22) is more singular when $|v| \approx 1$, i.e., for large w .

III. DISTRIBUTION OF b_z

The complex numbers b_z , that determine the asymptotics of the bound states of U , can be expressed in terms of the limiting behavior of the eigenfunctions. Specifically, for $\eta > 0$ we have from Eqs. (4) and (6),

$$\Psi_z(t) \rightarrow \begin{pmatrix} 0 \\ 1 \end{pmatrix} e^{izt}, \quad \Psi_z(-t) \rightarrow \begin{pmatrix} b_z^{-1} \\ 0 \end{pmatrix} e^{izt} \quad \text{as } t \rightarrow \infty. \quad (24)$$

Defining $\tilde{\Psi}_z(t) \equiv \Psi_z(-t)$ we can write

$$b = \lim_{t \rightarrow \infty} b(t), \quad b(t) \equiv \frac{\psi_2(t)}{\tilde{\psi}_1(t)}, \quad (25)$$

where for convenience we have dropped the subscript z . The time evolution of $\ln b$ is found from Eq. (2),

$$\frac{\partial \ln b}{\partial t} = i(uf + \tilde{u}^* \tilde{f}), \quad (26)$$

with $f(t) = \frac{\psi_1}{\psi_2}$, $\tilde{f}(t) = \frac{\tilde{\psi}_2}{\tilde{\psi}_1}$, $\tilde{u}(t) \equiv u(-t)$, and

$$\begin{aligned} \frac{\partial f}{\partial t} &= -2izf + iu^* - iuf^2, \\ \frac{\partial \tilde{f}}{\partial t} &= -2iz\tilde{f} - i\tilde{u} + i\tilde{u}^* \tilde{f}^2. \end{aligned} \quad (27)$$

A. Circular complex Gaussian $u(t)$

For a circularly symmetric complex Gaussian u , the solution of the Fokker-Planck equation derived from Eqs. (26) and (27) relaxes for large times toward a stationary solution where $\ln b$ is uniformly distributed [43], while f and \tilde{f} , expressed in polar form, i.e., $f = e^{w+i\phi}$, $\tilde{f} = e^{\tilde{w}+i\tilde{\phi}}$, are distributed independently according to the steady state solution (16). Because of the infinite range of the real part of $\ln b$ however, this stationary solution is ill defined. A better approach is to discretize the size T of the pulse into steps of size τ , equal to the inverse bandwidth of the input signal. Equation (26) then reads

$$\ln b = i\tau \sum_{i=1}^{T/2\tau} (u_i f_i + \tilde{u}_i^* \tilde{f}_i). \quad (28)$$

The variables u_i, \tilde{u}_i are *i.i.d.* Gaussian random variables, distributed according to

$$P_u(u) = \frac{1}{\sqrt{2\pi}u_0} e^{-u^2/2u_0^2}, \quad (29)$$

where $u_0^2 = D/2\tau$ and u stands for the real or imaginary part of either variable. For large enough T , the sum in Eq. (28) will be dominated by the domain where the distributions of f_i and \tilde{f}_i have reached their steady state. In this domain, we find that the real and imaginary parts of the products $x_i \equiv u_i f_i$ and $\tilde{x}_i \equiv \tilde{u}_i^* \tilde{f}_i$ have zero mean and the tails of their distributions fall off as the inverse third power of the argument. More precisely,

$$P(\chi) \sim 4\sqrt{\pi}u_0^2 \frac{\eta e^{2\eta/D}}{D \sinh(\frac{2\eta}{D})} \frac{1}{|\chi|^3}, \quad |\chi| \rightarrow \infty, \quad (30)$$

where χ stands for the real and imaginary parts of x_i, \tilde{x}_i . The general theory for sums of random variables [44–46] then tells us that for large T/τ the distribution of $\ln|b|$ will be Gaussian, with zero mean and variance

$$\sigma^2 = 4\sqrt{\pi} \frac{\eta e^{2\eta/D}}{\sinh(\frac{2\eta}{D})} T \ln \frac{T}{2\tau}. \quad (31)$$

The imaginary part of $\ln b$ is an angle and so, although it follows the same distribution as the real part, will, due to periodicity, become uniformly distributed in $[0, 2\pi)$. As seen in Eq. (6) for $\eta < 0$ the corresponding b_z is replaced by $-b_z^*$ [37]. Thus their distribution will be the same as that of the b 's, with η replaced by $-\eta$ in Eq. (31).

B. Real Gaussian $u(t)$

In this case, the Fokker-Planck equation for $\ln b$, derived from Eqs. (26) and (27) after setting u_2 and \tilde{u}_2 to zero, again predicts that its distribution becomes uniform as the duration T of the pulse grows to infinity. Equation (26) can be written as

$$\begin{aligned} \frac{\partial \ln b}{\partial t} &= -(ue^w \sin \phi + \tilde{u}e^{\tilde{w}} \sin \tilde{\phi}) \\ &+ i(ue^w \cos \phi + \tilde{u}e^{\tilde{w}} \cos \tilde{\phi}). \end{aligned} \quad (32)$$

As seen above, an exact solution to Eq. (19) is not available, but we can still obtain the first terms of an expansion of the stationary probability distribution in powers of e^{-w} as follows:

$$P(w, \phi) \approx \alpha_1 e^{-w} \delta(\cos \phi) + \alpha_2 e^{-2w} + \mathcal{O}(e^{-3w}), \quad (33)$$

with an identical expansion for the distribution of \tilde{w} and $\tilde{\phi}$. The constants α_1, α_2 depend on ξ/D and η/D , but, being related to the normalization, they cannot be determined without a knowledge of the full solution. We can thus only partially specify the manner in which the real and imaginary parts of $\ln b$ approach uniformity as T grows.

As in the complex case, the real part, $\ln|b|$, will be a sum of independent variables $\chi_i = u_i e^{w_i} \sin \phi_i$. However, in this case, due to the more singular behavior of $P(w, \phi)$ for large w , the tails of χ_i will be longer, falling off as $1/|\chi_i|^2$ for large T . As a result, the distribution of $\ln|b|$ will asymptotically follow a Cauchy distribution scaling like T/τ [45,46]. Its statistical median will be zero by symmetry, coming from the even parity of the Gaussian distribution of u, \tilde{u} . The phase of b does not get contributions from the first term in Eq. (33) because of the delta function in this term. For large T , the second term in the expansion dominates, making it uniform over $[0, 2\pi)$, in the same manner we saw in the case of complex u . Note that for the special case of $\xi=0$, the exact solution [cf. Eq. (22)] is proportional to $\delta(\cos \phi)$. The scale parameter of the Cauchy distribution will be

$$\gamma \sim \frac{e^{\eta/D}}{I_0(\eta/D)} \frac{T}{\tau}. \quad (34)$$

Only the transients of the distribution add to the phase of b , and numerical simulation shows that they are enough to again make it uniform.

IV. DISCUSSION

In the context of the NLSE, the scattering data of the ZS operator uniquely determine the solitonic excitations we get

in the optical fiber if we feed one end with a delta-correlated Gaussian signal. Even though the informational contents of the Gaussian signal and its solitonic spectrum are the same, it is easier to consider the effect of amplifier noise in the domain of the scattering data. For example, a small amount of amplifier noise will randomly shift each eigenvalue z by a small amount, while making large changes in the output signal [31]. The effect of this noise is important to analyze, in order to calculate the ultimate information capacity limits through optical fibers. In principle, to find the capacity one needs to optimize over input signal distributions, which is a formidable task. Instead, in this paper we start with a given input distribution and calculate the corresponding density of states and the corresponding distribution of scattering data b_z . We leave the analysis of the effects of noise on the spectrum for a future publication.

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APPENDIX: THOULESS FORMULA

The proof of the Thouless formula for the ZS eigenproblem proceeds similarly to the proofs in [25,29]. We consider the system of Eqs. (2) on the interval $[0, T]$. Let $\Psi^l(t), \Psi^r(t)$ be two independent solutions of Eqs. (2) that satisfy the conditions

$$\Psi^l(0) = \begin{pmatrix} \psi_1^l(0) \\ \psi_2^l(0) \end{pmatrix}, \quad \Psi^r(T) = \begin{pmatrix} \psi_1^r(T) \\ \psi_2^r(T) \end{pmatrix}. \quad (\text{A1})$$

We will need to combine this pair of initial and final conditions into a set of boundary conditions for the eigenstates, and for this we let each of them be a one-parameter family of initial (final) conditions to avoid overdetermining the problem. This means that $\psi_{1,2}^l(0)$ are not chosen independently, but satisfy a single linear relation. The same goes for $\psi_{1,2}^r(T)$. The Wronskian of the two solutions, $W = \psi_1^l \psi_2^r - \psi_2^l \psi_1^r$ is constant. Taking the derivative of Eq. (2) with respect to z , we obtain an equation for $\partial_z \Psi^l(t)$ whose solution can be written in terms of a matrix Green function \mathbf{G}^l as follows:

$$\frac{\partial}{\partial z} \Psi^l(t) = - \int_0^t dt' \mathbf{G}^l(t, t') \Psi^l(t'), \quad (\text{A2})$$

with

$$\mathbf{G}^l(t, t') = \begin{cases} \frac{i}{W} (\Psi^r(t) (\sigma_1 \Psi^l(t'))^T - \Psi^l(t) (\sigma_1 \Psi^r(t'))^T), & t > t' \\ 0, & t < t'. \end{cases} \quad (\text{A3})$$

Here $\sigma_1 = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}$. The matrix Green function \mathbf{G}^l satisfies the initial conditions

$$\mathbf{G}^l(0, t') = 0, \quad \left. \frac{\partial}{\partial t} \mathbf{G}^l(t, t') \right|_{t=0} = 0. \quad (\text{A4})$$

We also define another Green function,

$$\mathbf{G}(t, t') = \begin{cases} \frac{i}{W} \Psi^r(t) (\sigma_1 \Psi^l(t'))^T, & t > t' \\ \frac{i}{W} \Psi^l(t) (\sigma_1 \Psi^r(t'))^T, & t < t', \end{cases} \quad (\text{A5})$$

which satisfies the conditions (A1) (taken together as boundary conditions) and will determine the density of states. The Lyapunov exponent can be expressed as [47]

$$\kappa = \lim_{T \rightarrow \infty} \frac{1}{2T} \langle \ln(\Psi^{l\dagger}(T) \Psi^l(T)) \rangle. \quad (\text{A6})$$

Before going any further, we must note that the value of κ is, with probability one, independent of the initial conditions satisfied by Ψ^l (the argument is very similar to that for the FGM [27]). To see this, we rewrite the system of Eqs. (2) as

$$i \frac{\partial}{\partial t} \begin{pmatrix} \psi_1 \\ \psi_2 \end{pmatrix} = \mathbf{V} \begin{pmatrix} \psi_1 \\ \psi_2 \end{pmatrix}, \quad \mathbf{V} = \begin{pmatrix} z & -u^* \\ -u & -z \end{pmatrix}. \quad (\text{A7})$$

We can formally write the solution for Ψ^l as follows:

$$\Psi^l(t) = \mathbf{S}(t, 0) \Psi^l(0), \quad (\text{A8})$$

$$\mathbf{S}(t, 0) = \mathcal{T} \exp \left[-i \int_0^t dt' \mathbf{V}(t') \right], \quad (\text{A9})$$

where “ $\mathcal{T} \exp$ ” denotes the path-ordered exponential [48]. Because the trace of \mathbf{V} vanishes, we have $\det \mathbf{S} = 1$ [49]. Therefore, if we denote the two eigenvalues of $\mathbf{S}(t, 0)$ by $s_{\pm}(t)$, with $|s_+(t)| > |s_-(t)| = |s_+(t)|^{-1}$, we have

$$\kappa = \lim_{T \rightarrow \infty} \frac{1}{2T} \ln(|s_+(T) \tilde{\psi}_+(T)|^2 + |s_-(T) \tilde{\psi}_-(T)|^2) = \lim_{T \rightarrow \infty} \frac{1}{T} \ln |s_+(T)|, \quad (\text{A10})$$

independent of the initial condition, given that the coefficient of the exponentially increasing solution does not vanish, which is the case with probability one in the limit of large T .

Now, since $\Psi^{l\dagger}(T)$ depends only on z^* and not on z , taking the derivative of Eq. (A6) with respect to the latter, we get

$$\frac{\partial \kappa}{\partial z} = \lim_{T \rightarrow \infty} \frac{1}{2T} \left\langle \frac{\Psi^{l\dagger}(T) \frac{\partial \Psi^l(T)}{\partial z}}{\Psi^{l\dagger}(T) \Psi^l(T)} \right\rangle. \quad (\text{A11})$$

The quantity inside the average can be computed from Eq. (A2) as follows:

$$\frac{\Psi^{l\dagger}(T)\partial_z\Psi^l(T)}{\Psi^{l\dagger}(T)\Psi^l(T)} = -i \frac{\Psi^{l\dagger}(T)\Psi^r(T)}{W} \frac{\int_0^T dt' (2\psi_1^l(t')\psi_2^l(t'))}{\Psi^{l\dagger}(T)\Psi^l(T)} + \int_0^T dt' \text{Tr } \mathbf{G}(t',t'). \quad (\text{A12})$$

The first term on the right-hand side of Eq. (A12) is almost surely $\mathcal{O}(1)$ in the limit of large T and so does not contribute to the average. We are thus left with

$$\frac{\partial\kappa}{\partial z} = \lim_{T \rightarrow \infty} \frac{1}{2T} \left\langle \int_0^T dt' \text{Tr } \mathbf{G}(t',t') \right\rangle. \quad (\text{A13})$$

Taking the derivative of Eq. (A13) with respect to z^* and using the relation [12] $\rho(z, z^*) = \frac{1}{\pi T} \frac{\partial}{\partial z^*} \langle \text{Tr } \mathbf{G} \rangle$ we arrive at the Thouless formula for the density of states as follows:

$$\rho(z, z^*) = \frac{2}{\pi} \frac{\partial^2 \kappa}{\partial z^* \partial z} = \frac{1}{2\pi} \left(\frac{\partial^2}{\partial \xi^2} + \frac{\partial^2}{\partial \eta^2} \right) \kappa. \quad (\text{A14})$$

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